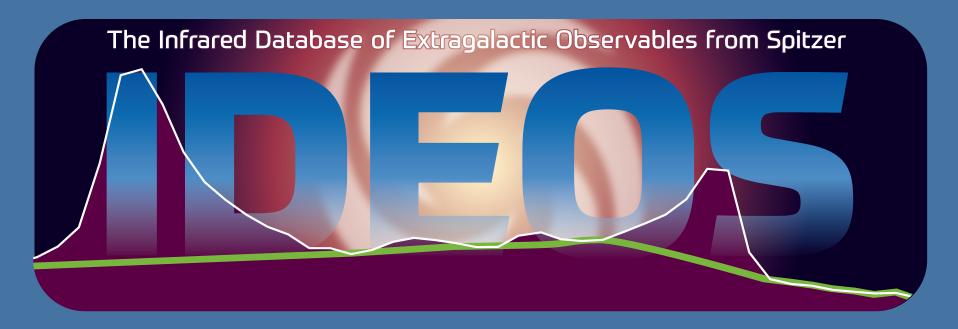


#### **IDEOS collaborators:**

Antonio Hernán-Caballero Vianney Lebouteiller Tom Loredo John Miles David Rupke

NASA-ADAP grants 2012 & 2015



- Observables for 3500 IRS low-resolution galaxy spectra:
- Line fluxes for 18 lines
- Fluxes and equivalent widths for 14 PAH bands
- 9.7µm silicate strength
- crystalline silicate strengths for 4 bands
- 6µm water ice; 6.85µm hydrocarbon band
- Synthetic photometry for 15 Spitzer, WISE and JWST-MIRI bands



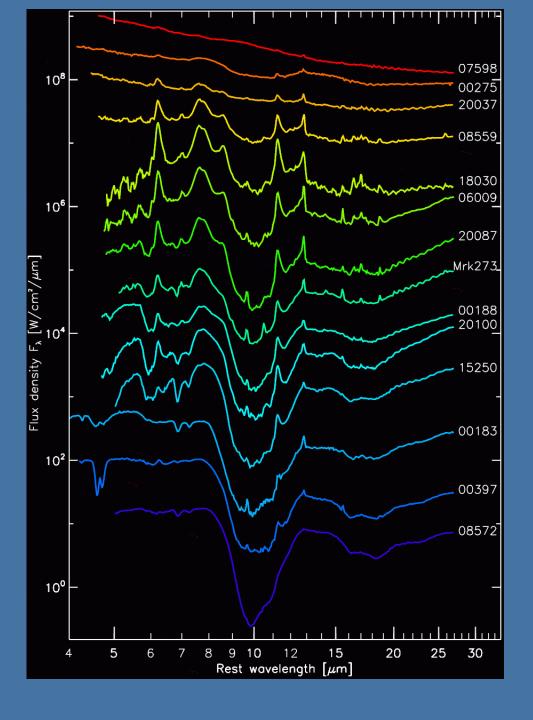
# 13,500 CASSIS spectra $\rightarrow$ 3,500 IDEOS spectra:

- select all *galaxy* spectra from CASSIS
- discard poor S/N spectra
- choose optimal or tapered extraction based on source size
- match IRS source to NED source within IRS error circle
- obtain spectroscopic redshift: zNED or zIRS. (124 galaxies had wrong zNED. 228 got their first redshift: zIRS)
- combine spectral segments from different observations
- Scale and stitch spectral segments

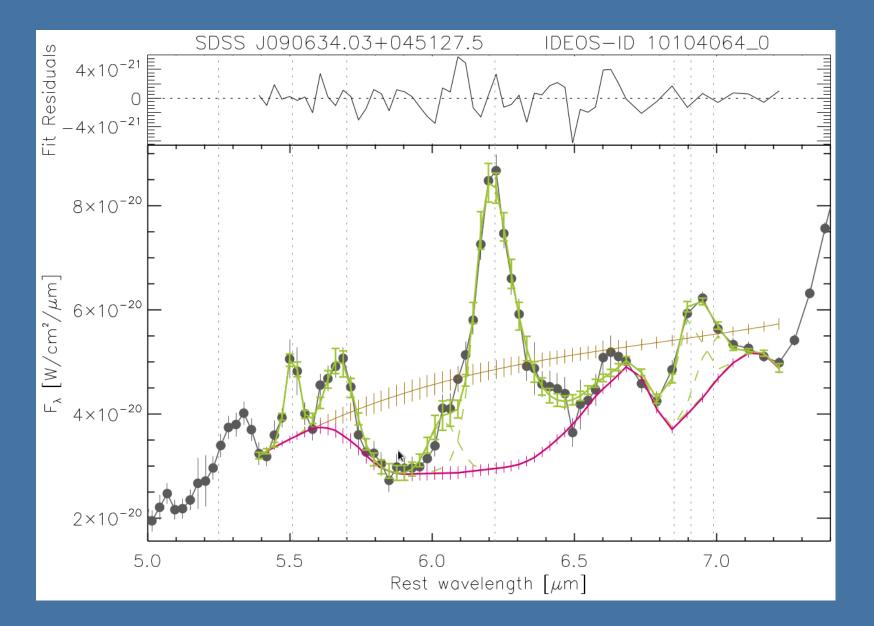
## Challenge:

Fit model should accommodate many different features that:

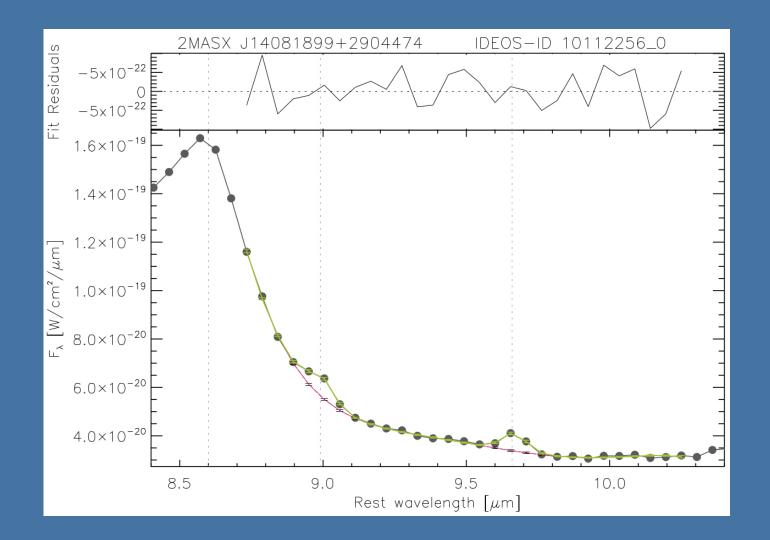
- overlap
- appear only in a fraction of sources
- Continuum needs to be very accurate to measure lines!



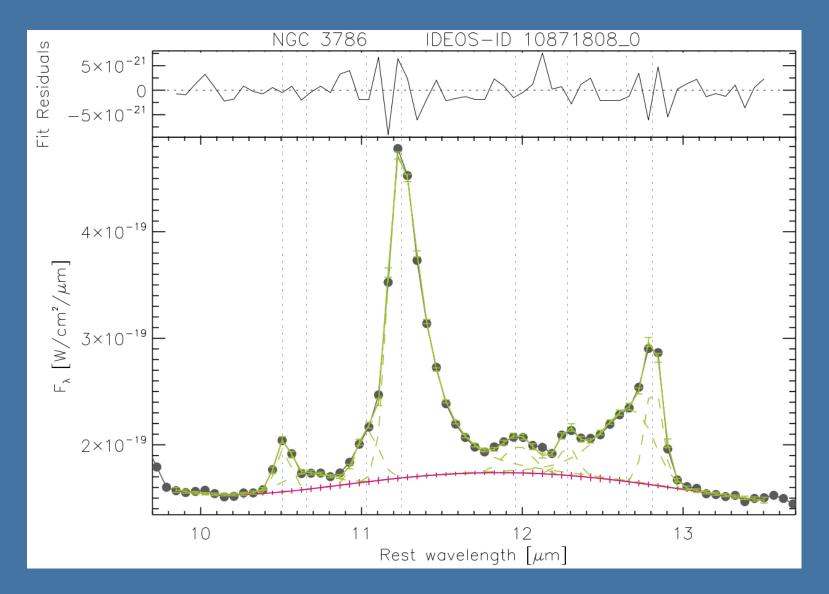
#### MPFIT model of 5.4-7.2µm range



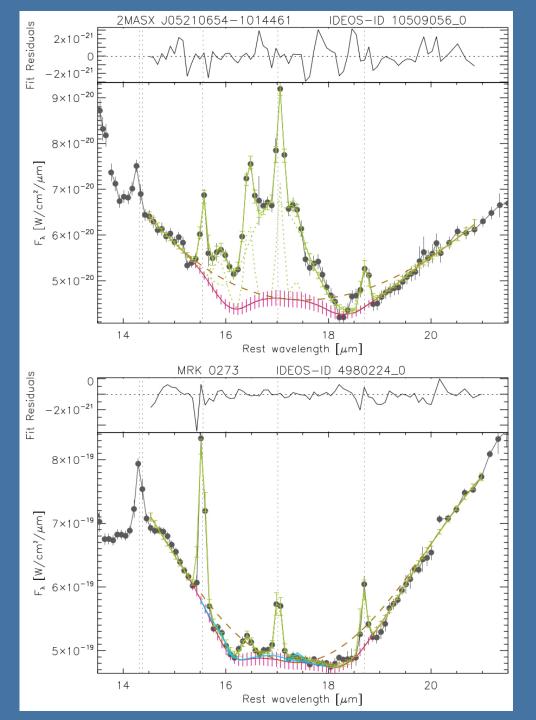
## 8.6-10µm range



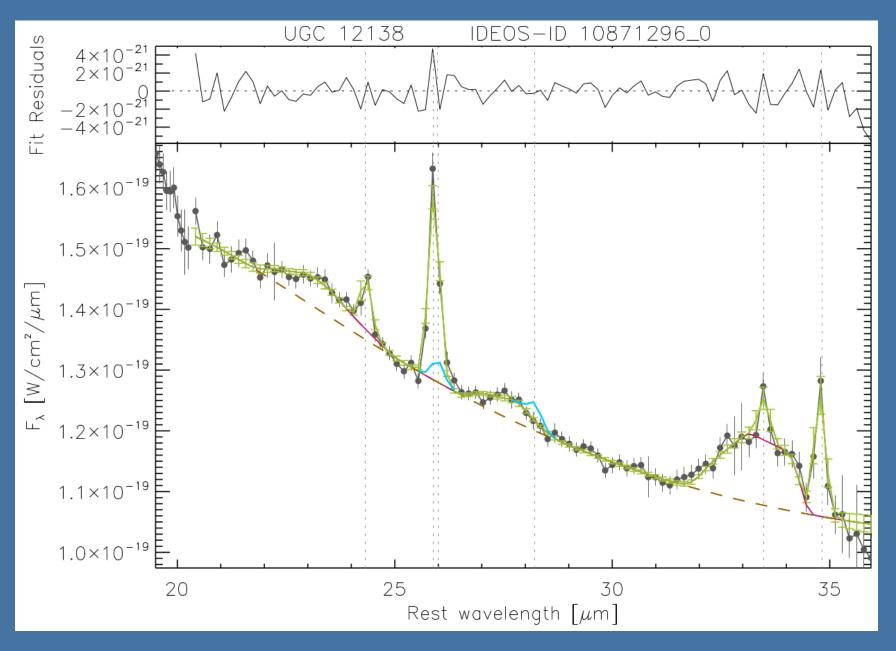
# 9.8-13.5µm range fit



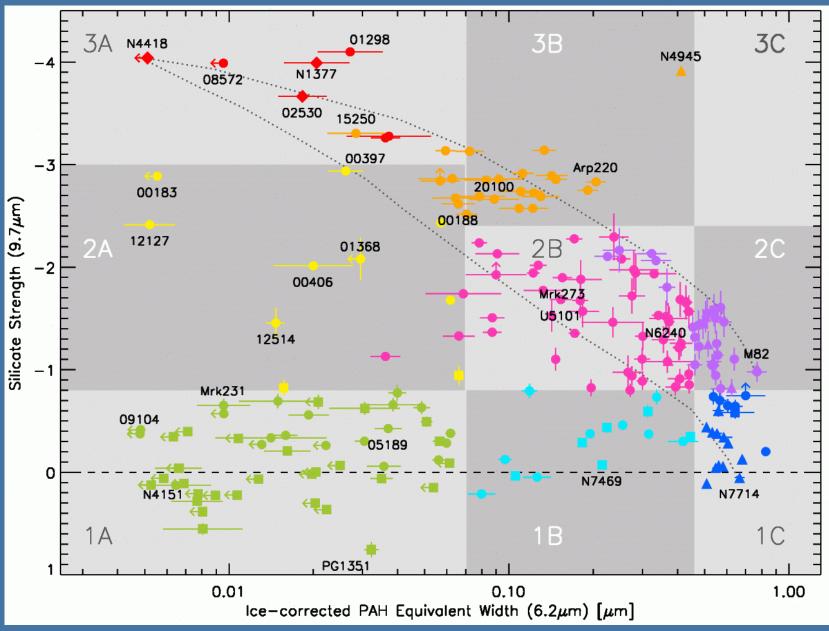
# 15-20µm range fit



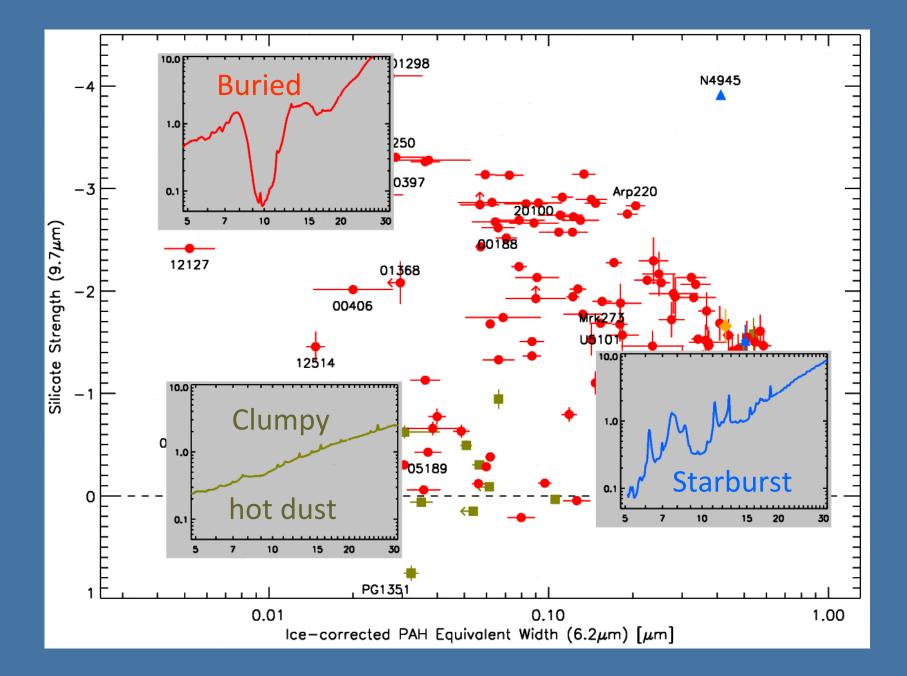
## 20-36µm range fit



#### Mid-IR spectral classification

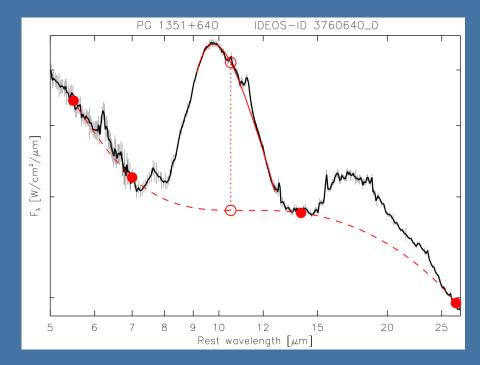


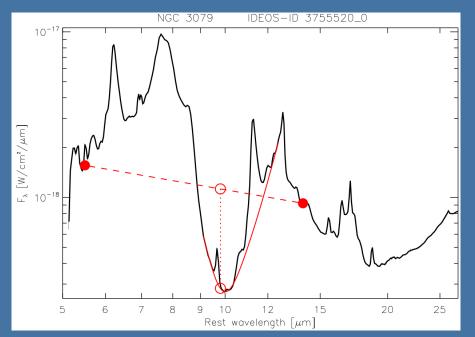
Spoon+07

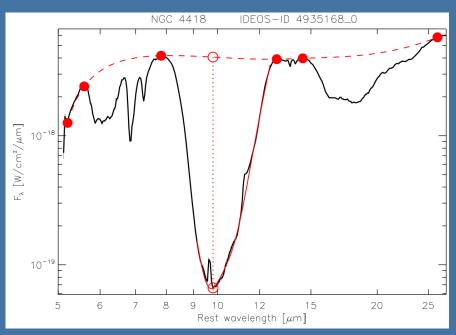


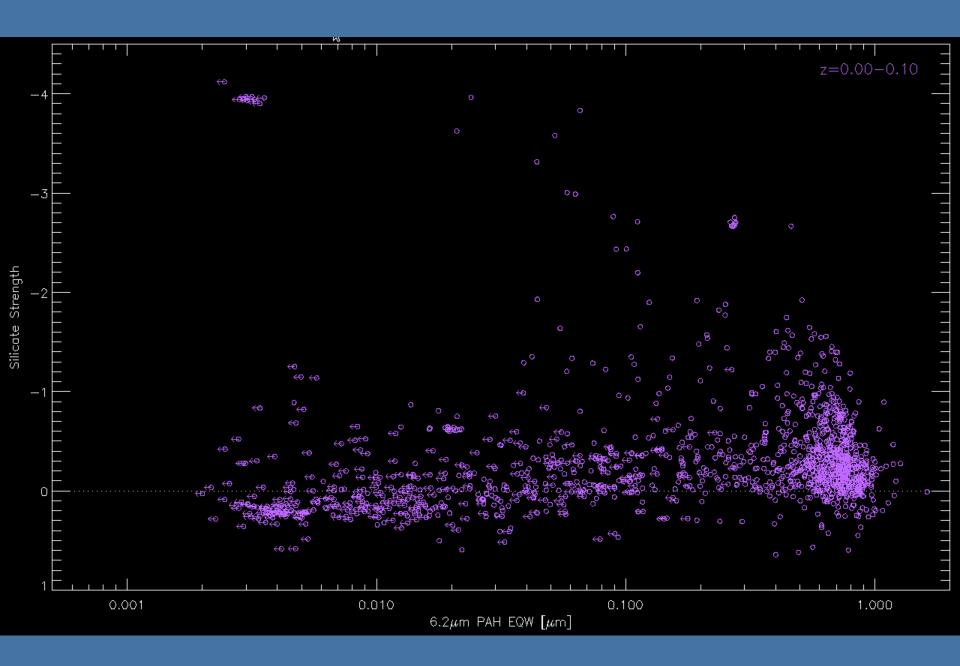
#### Determining Silicate Strength S<sub>sil</sub>:

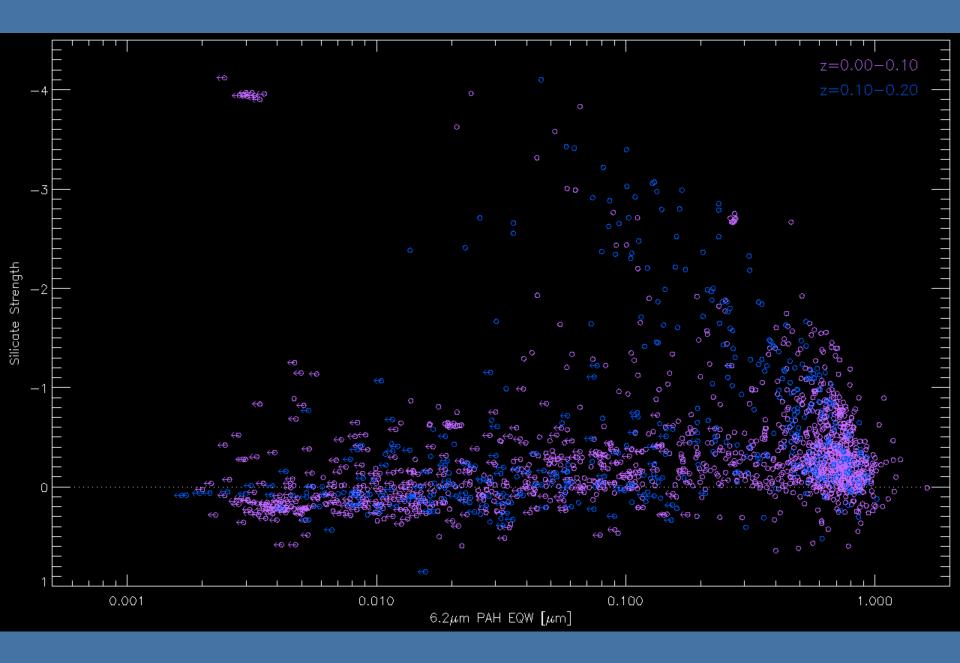
- $S_{sil} = LN [Cont_{local} / Cont_{interpolated}]$ at 9.8µm or 10.5µm
- fit the local observed continuum
- interpolate over the local continuum:
  - power-law interpolation
  - spline interpolation

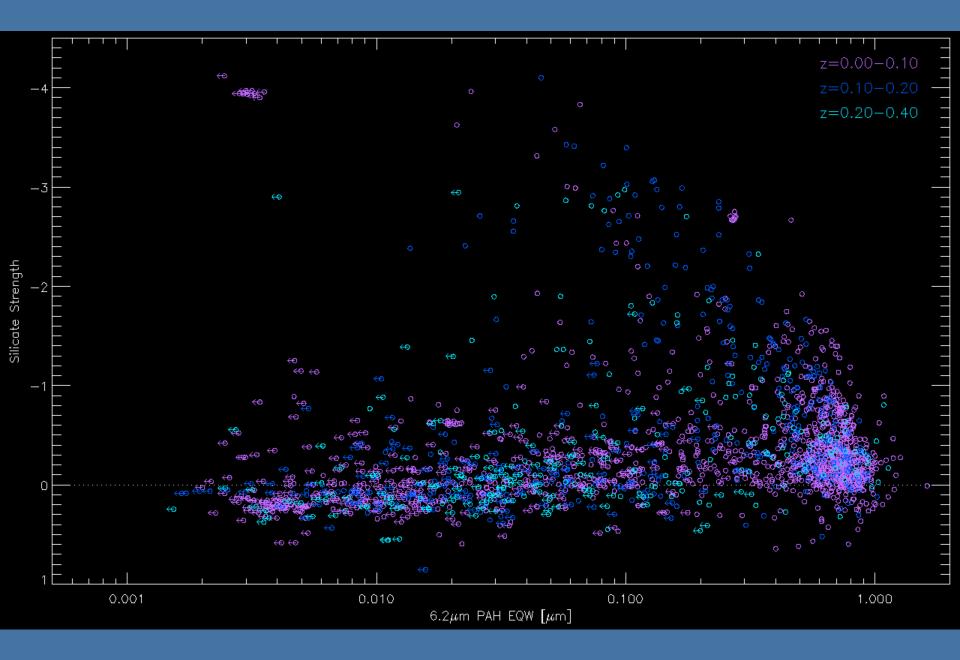


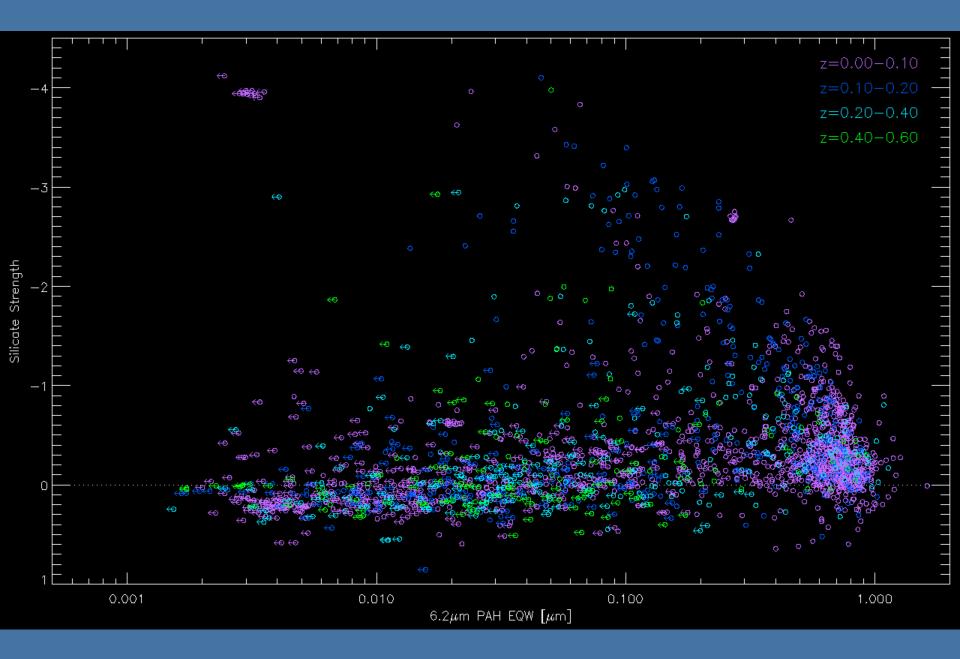


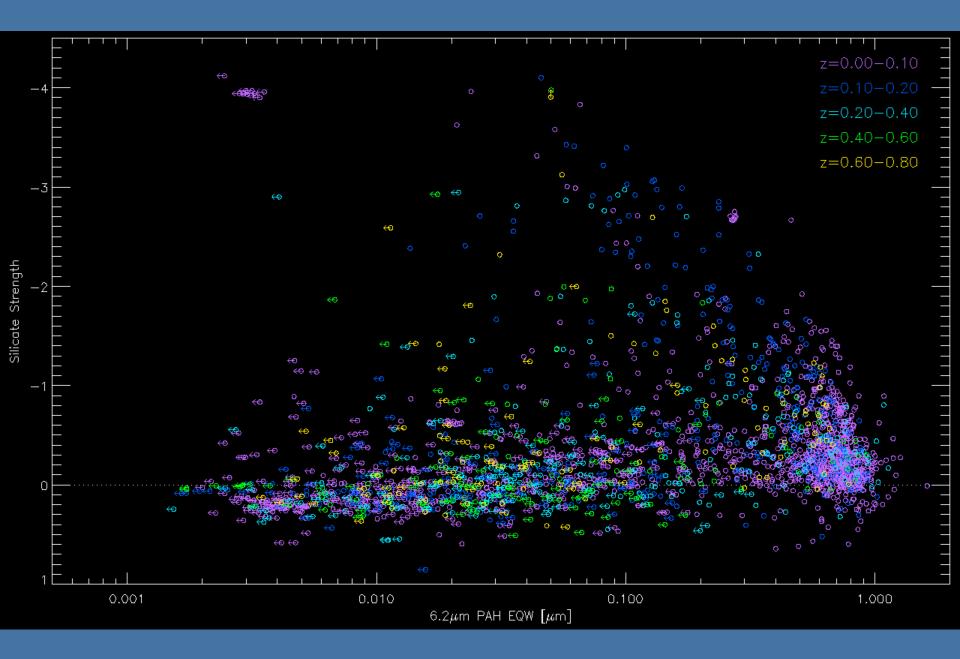


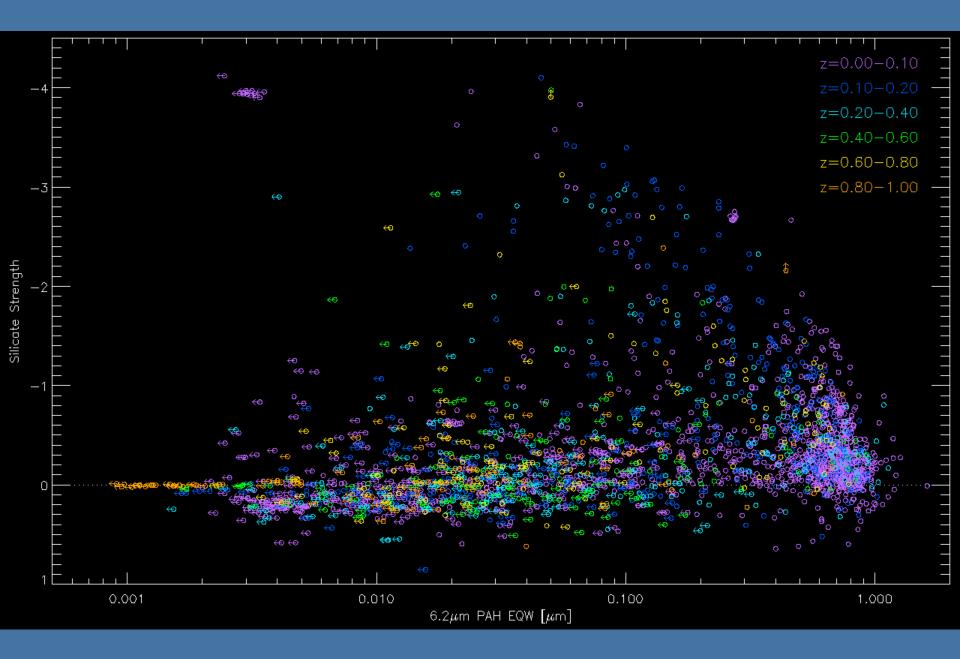


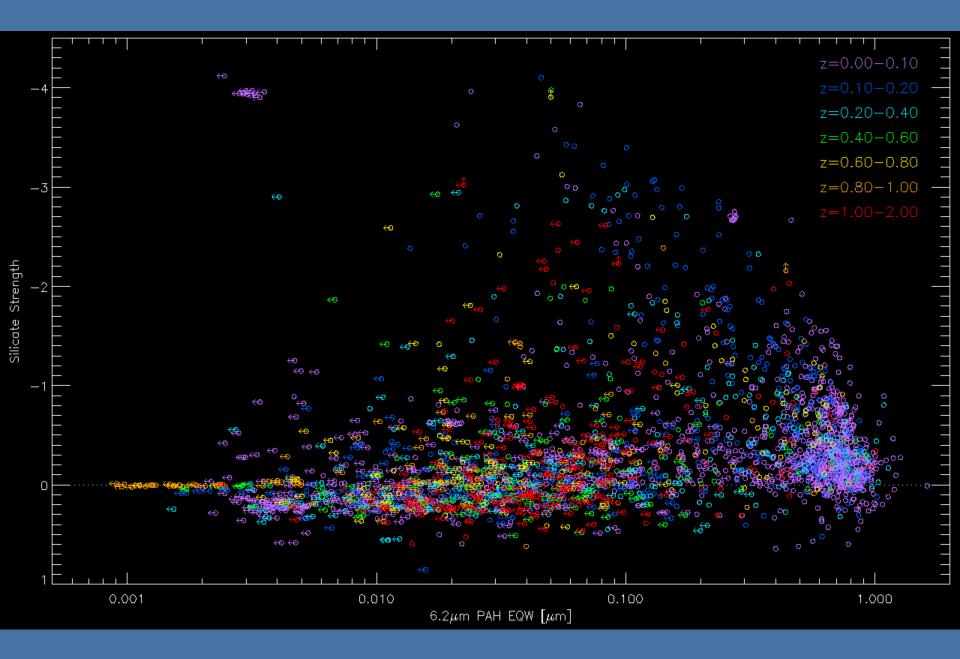




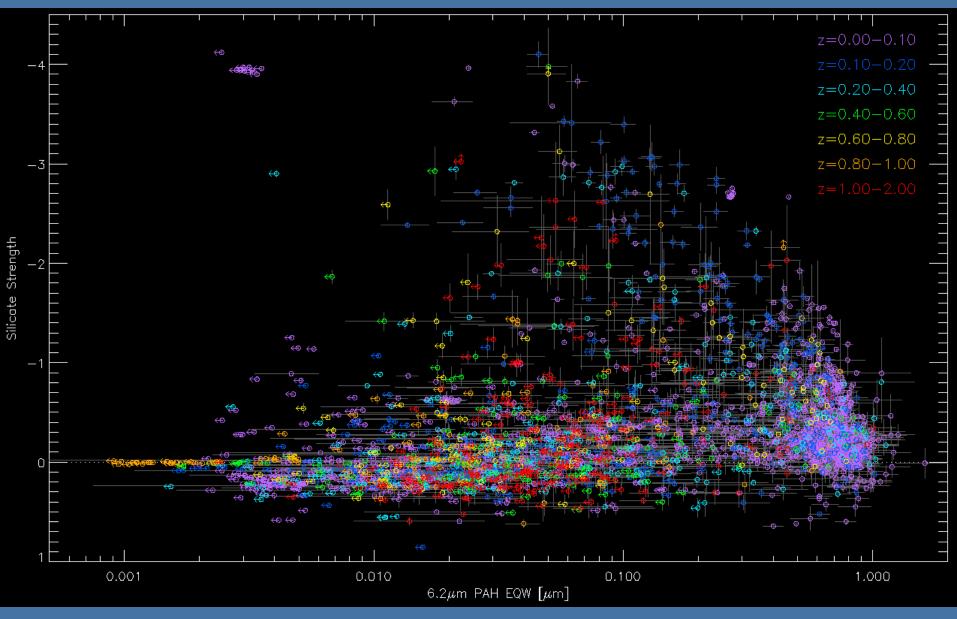




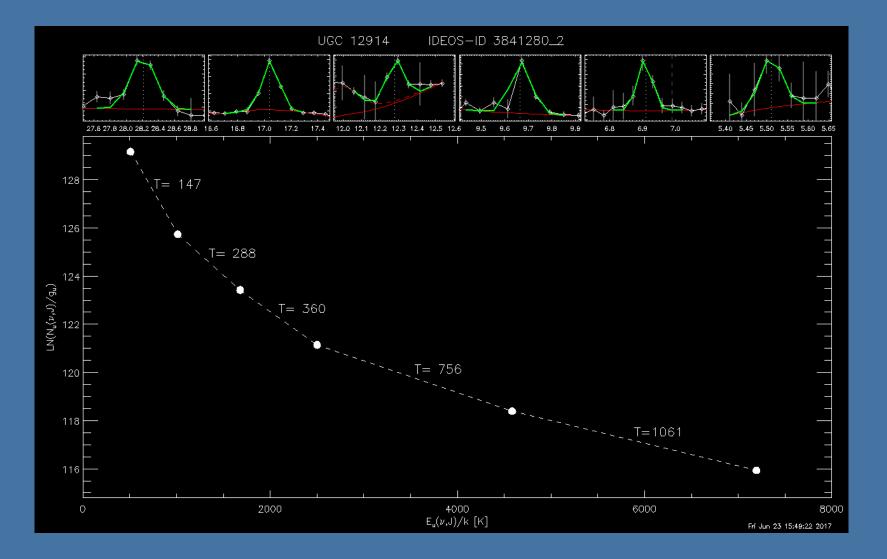




2158 IDEOS galaxies

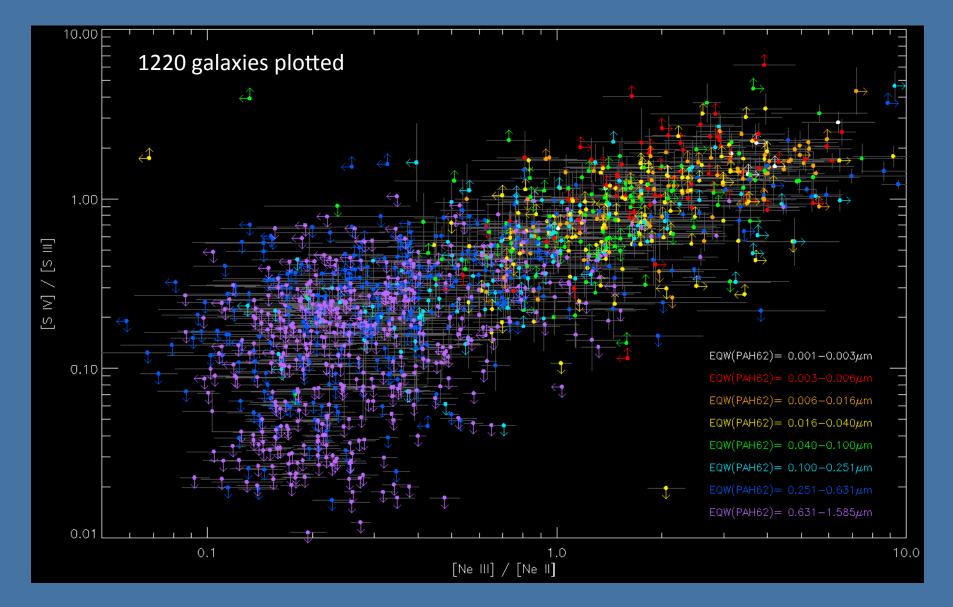


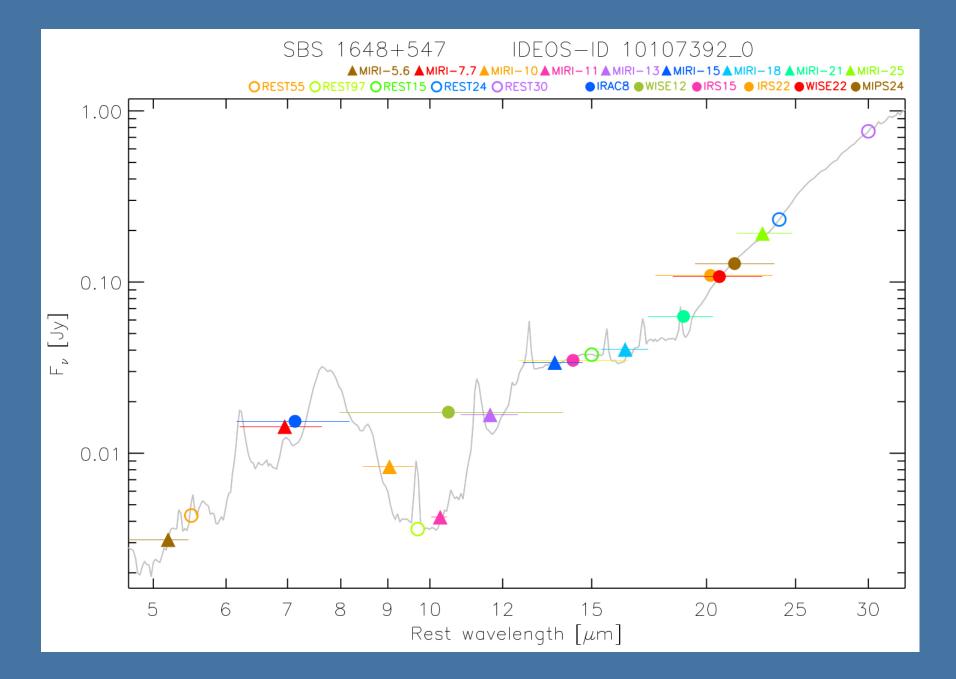
# H<sub>2</sub> excitation diagram



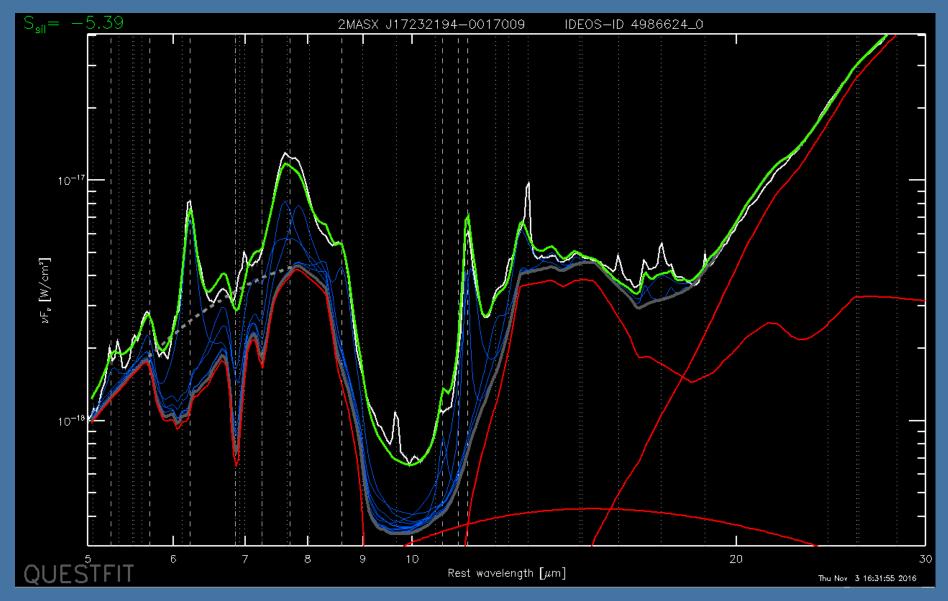
NGC 6764 IDEOS-ID 3858176\_0  $H_2 S(0)$  $H_2 S(1)$  $H_2 S(3)$  $H_2 S(2)$ 27.6 27.8 28.0 28.2 28.4 28.6 28.8 16.6 16.8 12.0 12.1 12.2 12.3 12.4 12.5 12.6 9.5 9.6 9.8 9.9  $H_2 S(5)$ [S Ⅲ]18.7µm [Ne II]  $H_2 S(7)$ 6.9 7.0 19.212.5 12.6 12.7 12.8 12.9 13.0 13.1 6.8 5.40 5.45 5.50 5.55 5.60 5.65 18.4 18.6 18.8 19.0 [Ar II] [Si II] [S Ⅲ]33.5µm [Ne III] 6.8 6.9 34.0 34.5 33.0 34.0 15.2 15.6 15.8 35.5 [Ar III] [O Ⅳ] & [Fe Ⅱ] [S Ⅳ] [Ne V]24.3µm 23.8 24.0 24.2 24.4 24.6 24.8 Tue Jun 20 12:57:57 2017 9.2 25.2 25.4 25.6 25.8 26.0 26.2 26.4 8.8 8.9 10.6

## **Diagnostic line ratios**





## PAHFIT and QUESTFIT spectral decomposition



QUESTFIT: Veilleux+09 PAHFIT: Smith+07

Crystallinity of Galactic ISM is low:

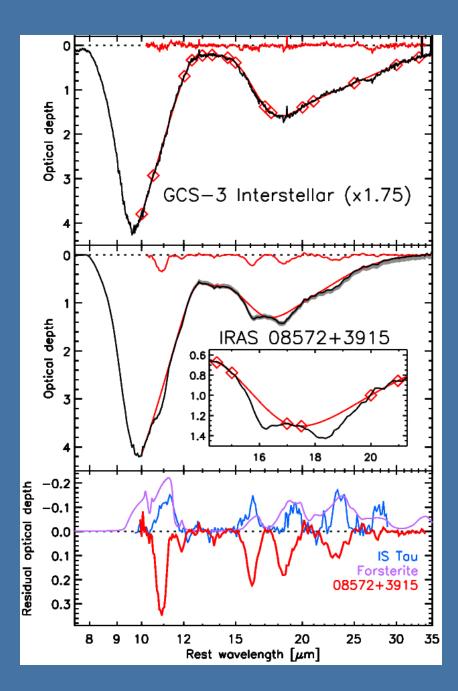
 $N_{cryst} / (N_{cryst} + N_{amor}) < 1\%$  (Kemper+04)

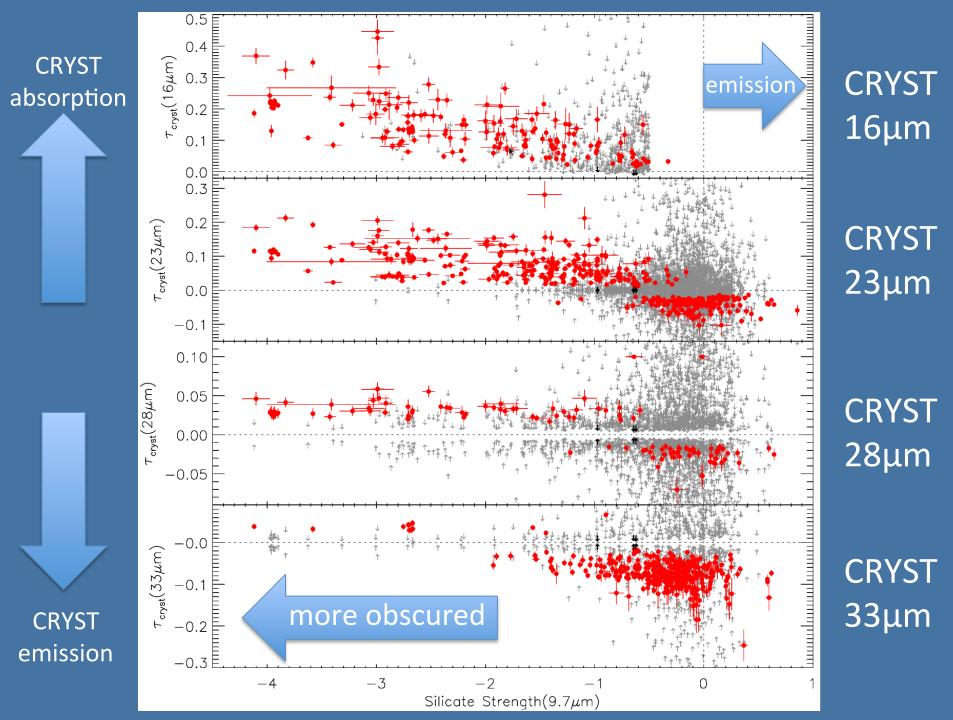
#### Crystallinity in ULIRG sample:

7-14% (Spoon+06)

#### IDEOS sample:

- we detected crystalline silicates in 844/3500 spectra
- strict non-detections in only 3/3500 spectra !

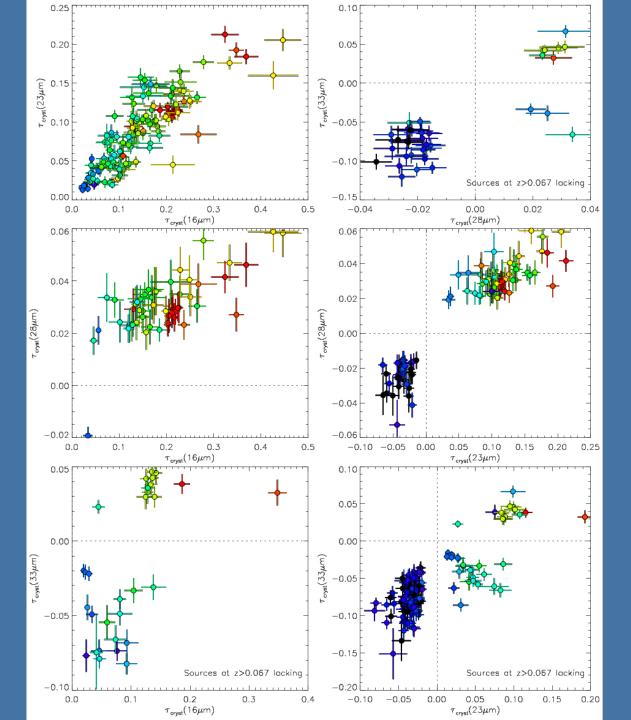




#### CRYST 16 & 23

#### CRYST 16 & 28

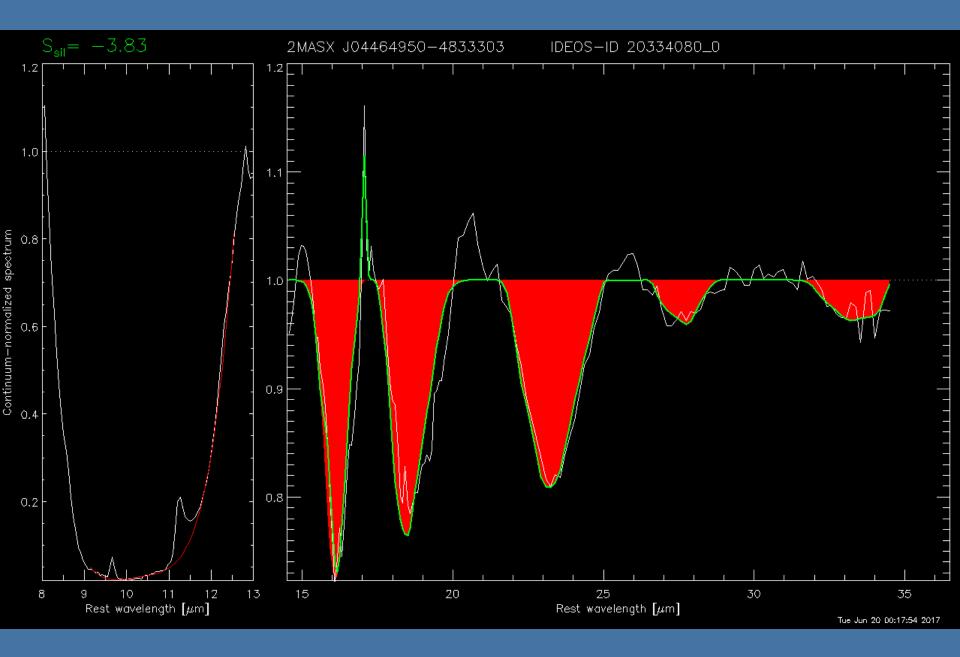
#### CRYST 16 & 33



## CRYST 28 & 33

#### CRYST 23 & 28

CRYST 23 & 33





# ideos.astro.cornell.edu

to become available soon: ask for a demo!

paper I:Hernán-Caballero+16 (redshifts) paper II: to be submitted end of summer'17 (database)